

Group 4

2012 ISR Summer School

Arpita Barua¹, Siddhartha Choudhury¹
Nathaniel A. Frissell², Vaibhav Kumar³,
Mandana Sobhanzadeh⁴

¹University of Saskatchewan

²Virginia Tech

³Stanford University

⁴University of Calgary

Mandatory Canadian Content



Outline

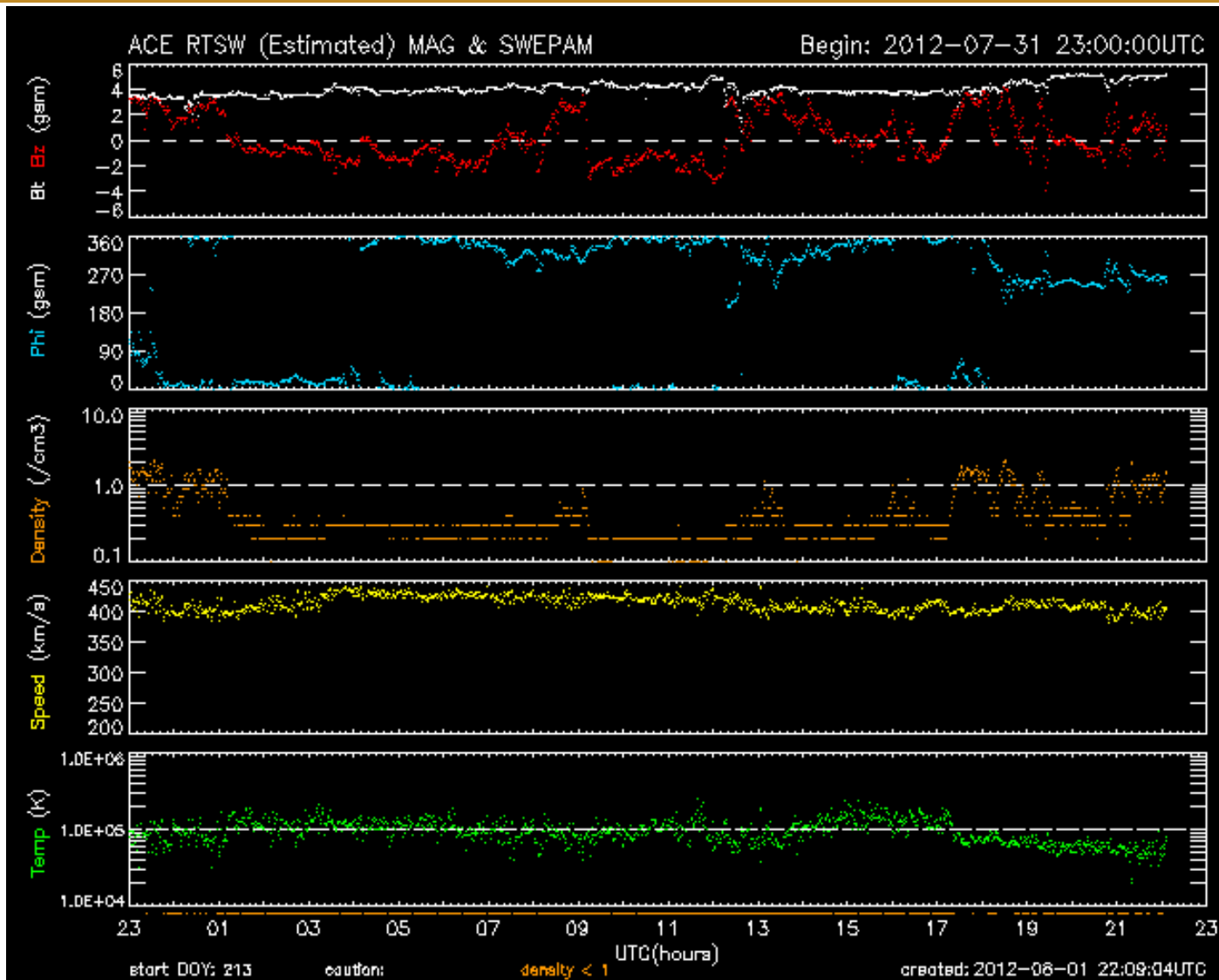
- I. Introduction
- II. Case Study: 1 August 2012 0000 – 1100 UT
 - I. Geospace and Geomagnetic Conditions
 - II. Poker Flat ISR (PFISR) Observations
 - III. Sondrestrom ISR Observations
- III. Summary and Conclusions

Introduction

- The auroral oval is region of particle precipitation typically located between $\sim 60 - 70^\circ$ magnetic latitude.
- Auroral oval exists even during quiet times.
- Dramatic nightside auroral brightening and enhanced precipitation occurs during substorms [*Akasofu, 1964*].
- Auroral observation provides a method of sensing magnetospheric dynamics and its impacts on the Earth.
- **Objective:** Characterize auroral structure and processes using ground-based instrumentation.

ACE SW/IMF

Figure courtesy of E. Cousins (Group 3)



Bz

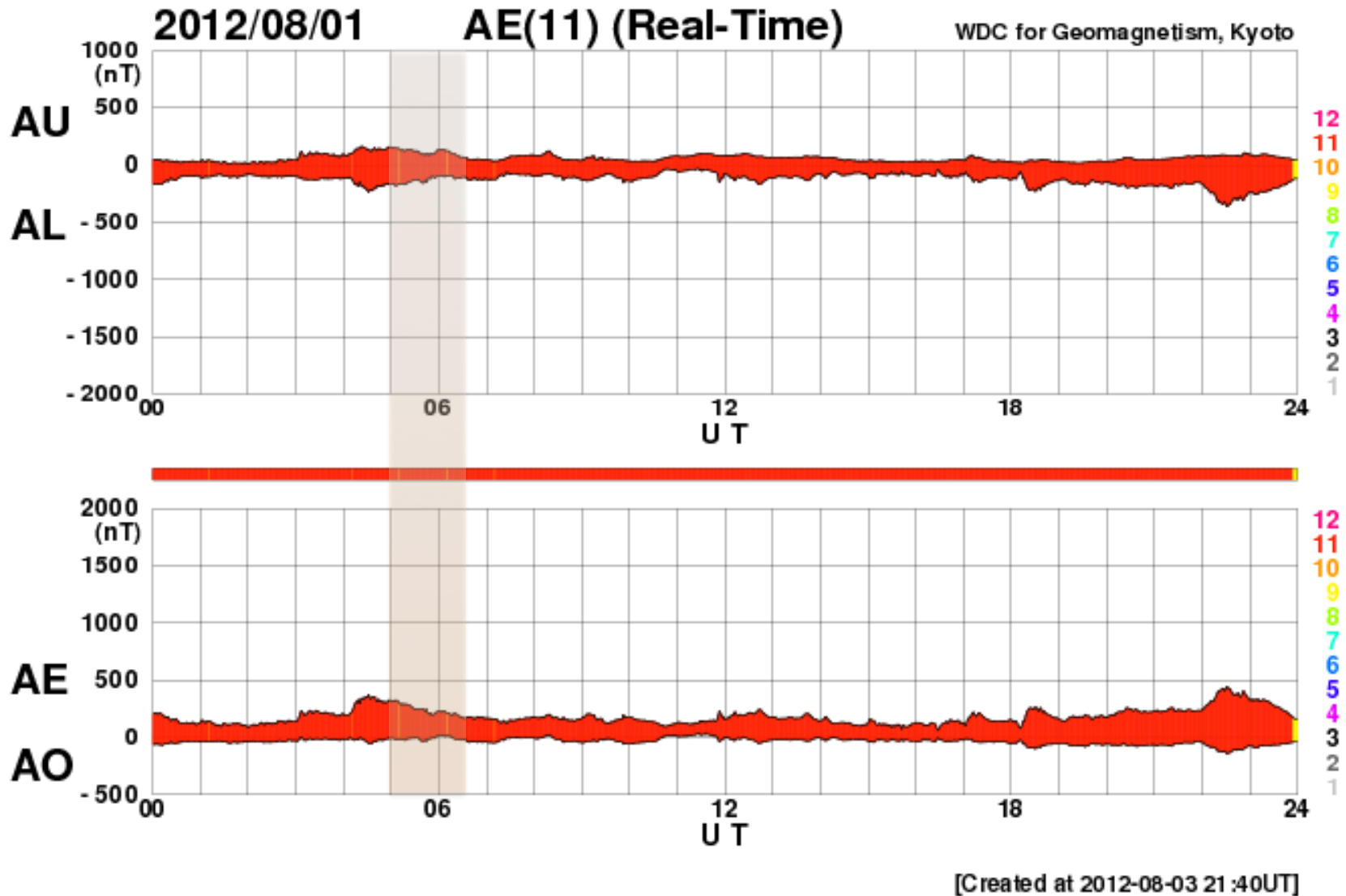
Phi

Density

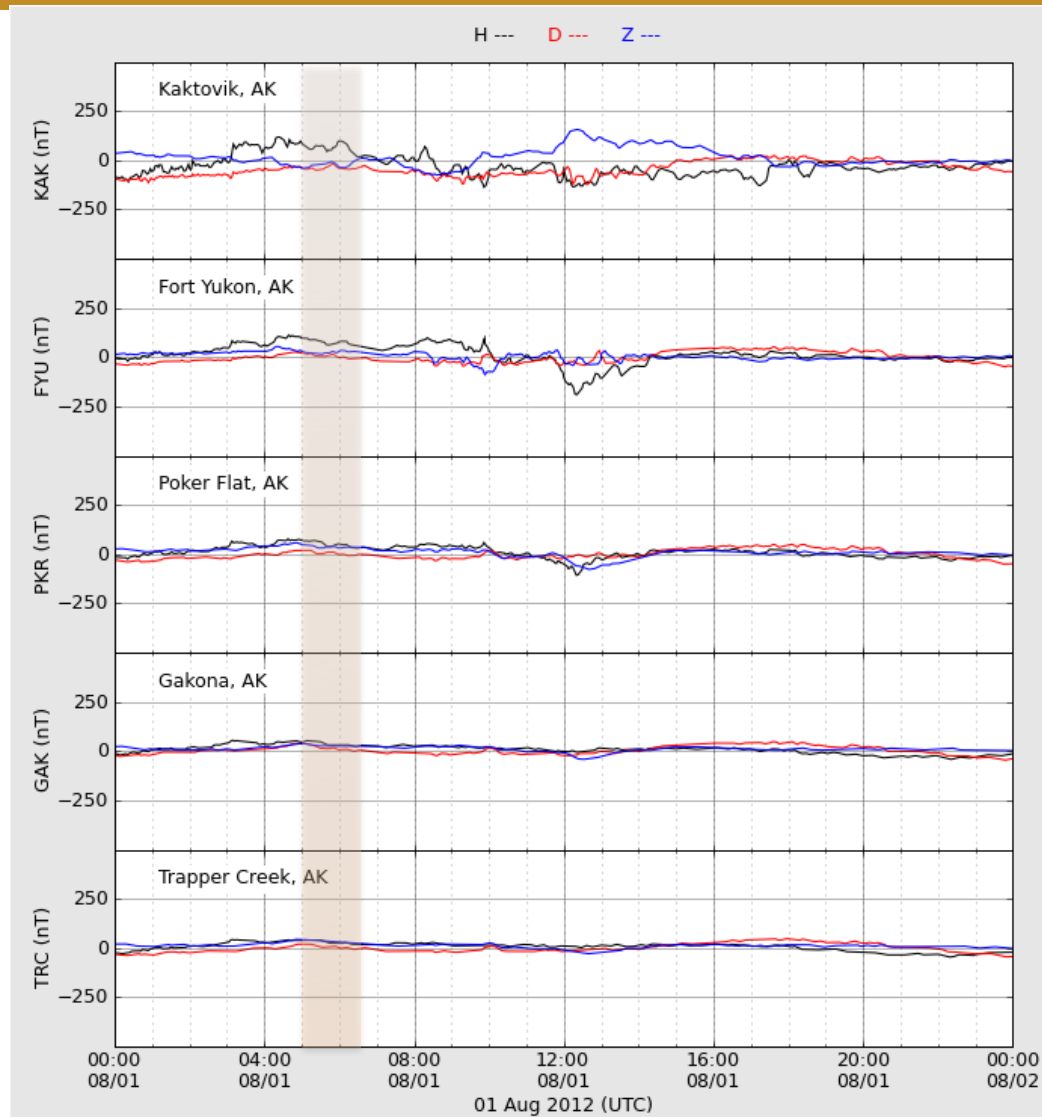
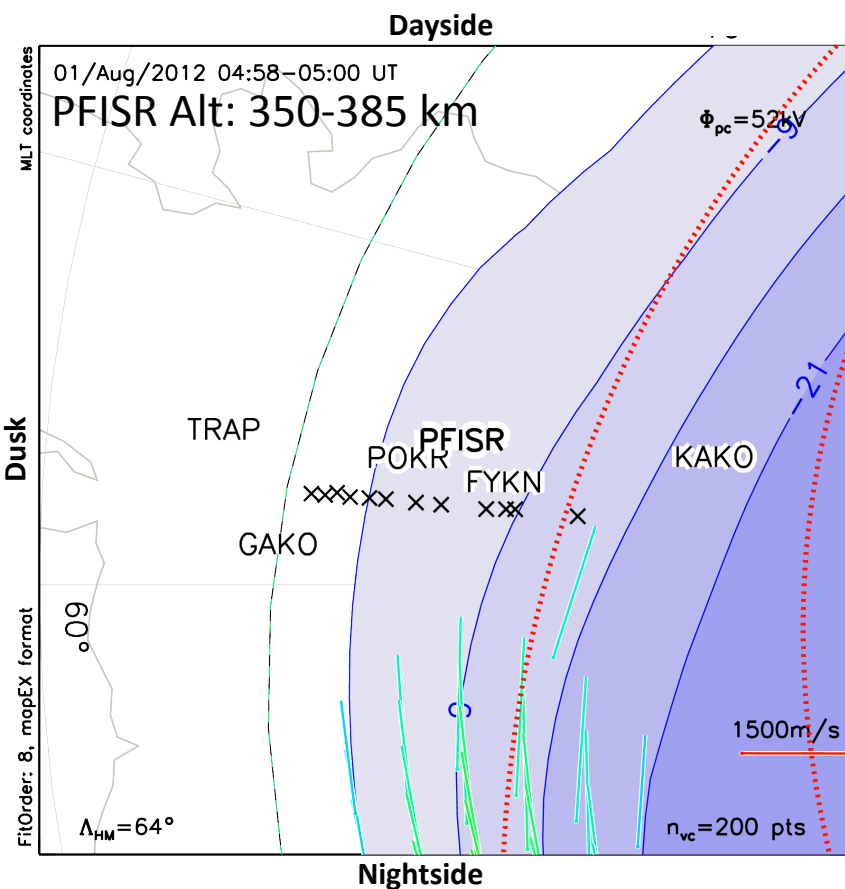
Speed

Temp

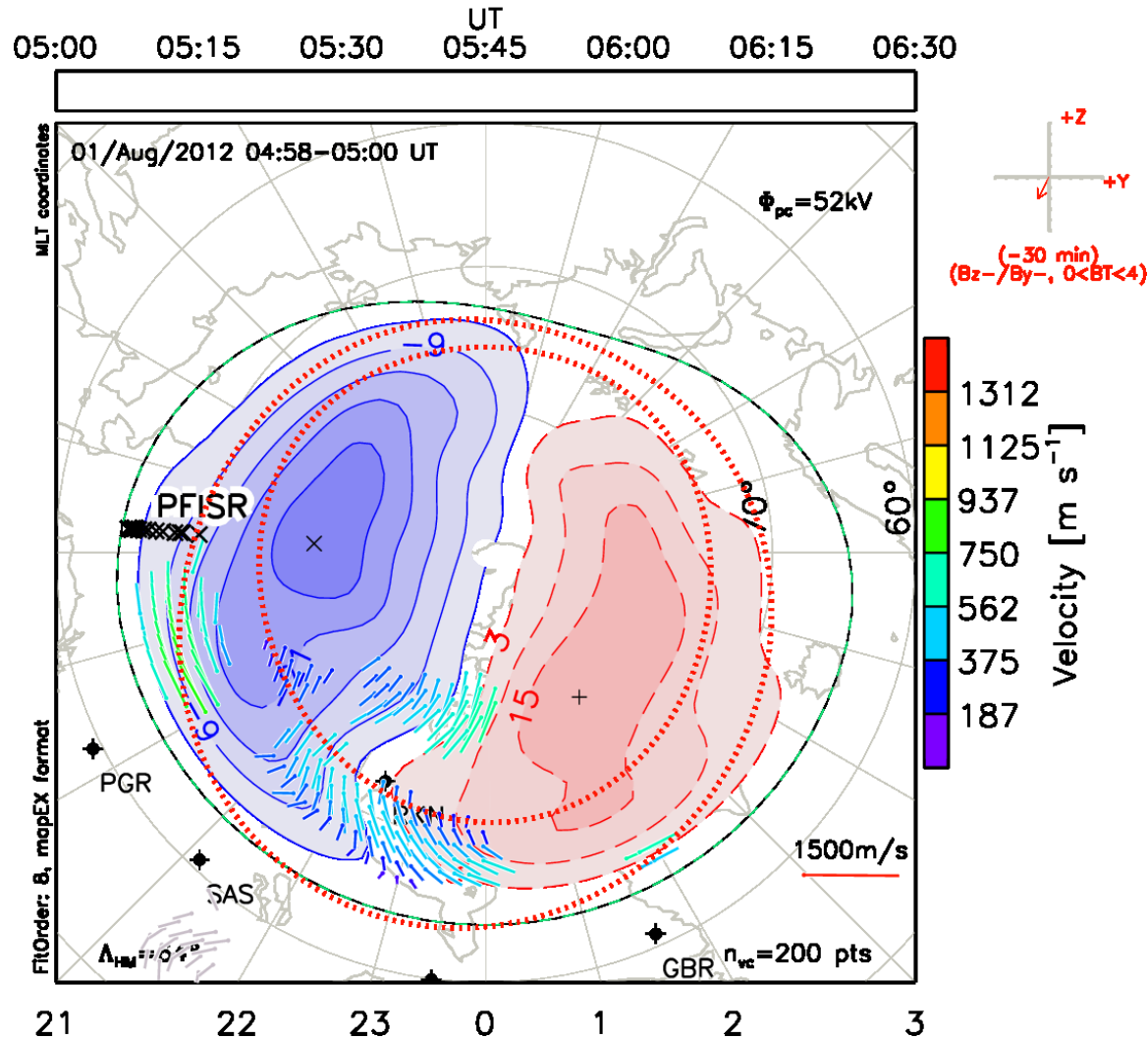
Geomagnetic Conditions



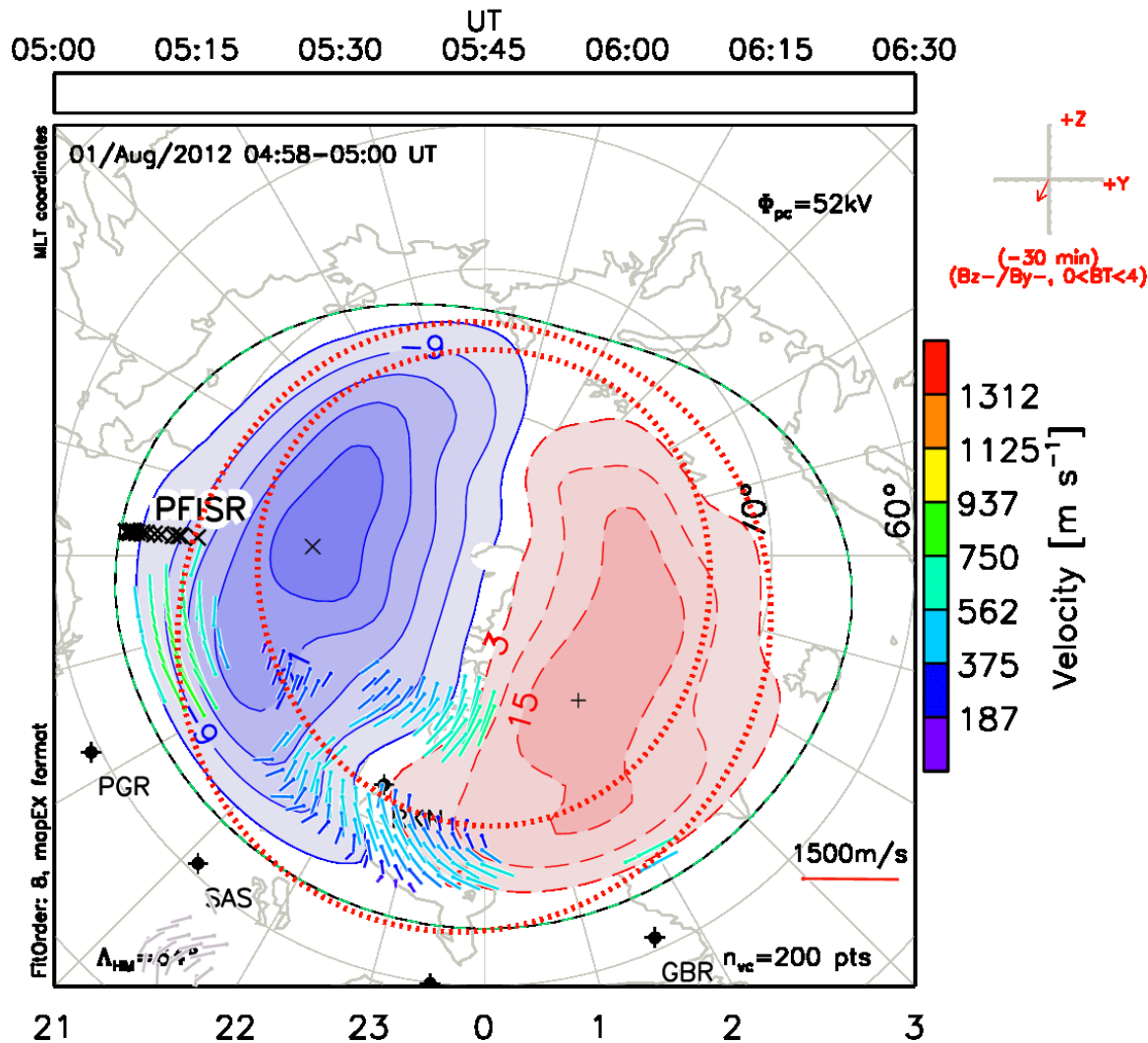
Ground Magnetometers



PFISR Altitude: 350–385 km; Felstein Oval for $K_p=1+$

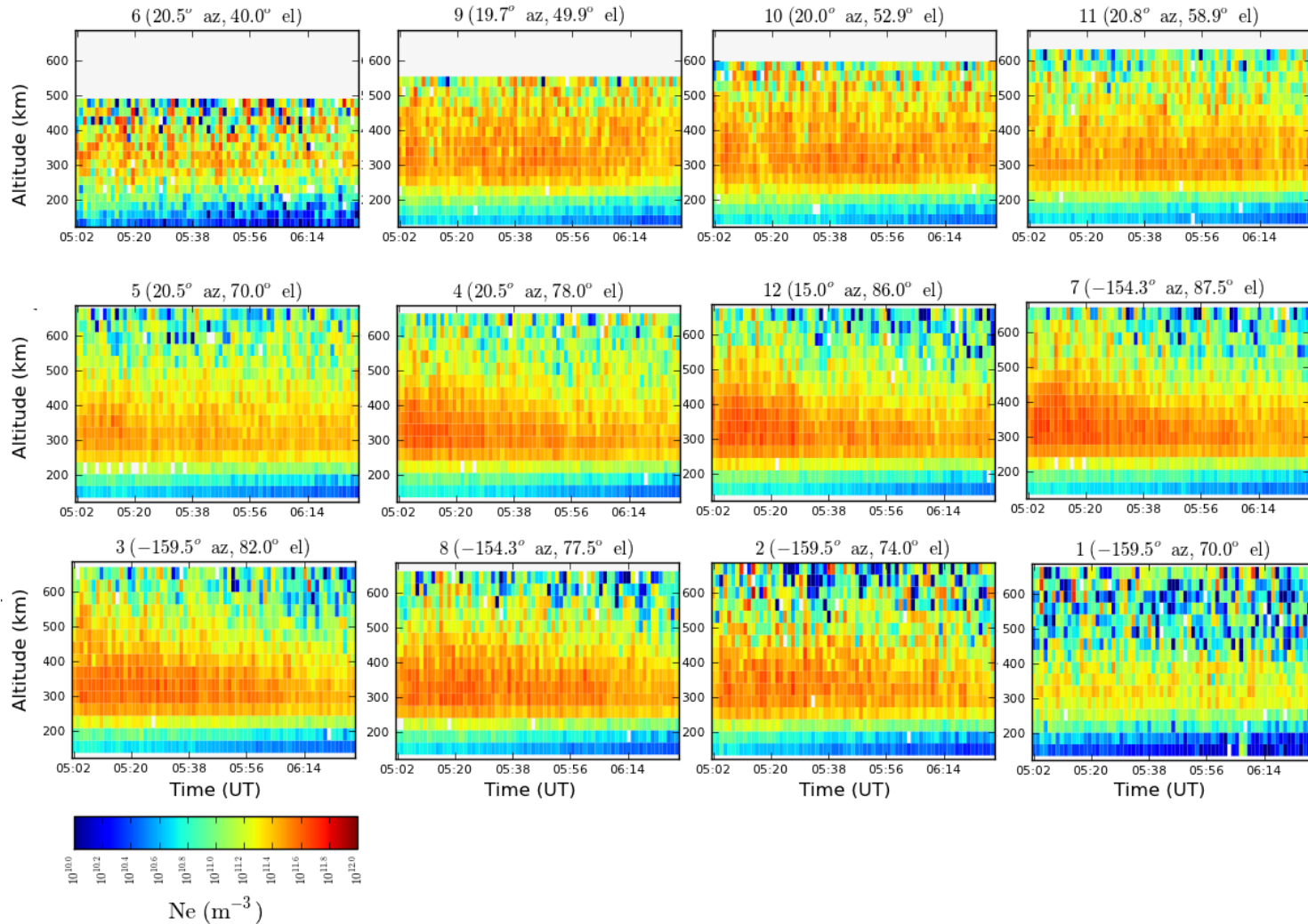


PFISR Altitude: 350–385 km; Felstein Oval for $K_p=1+$



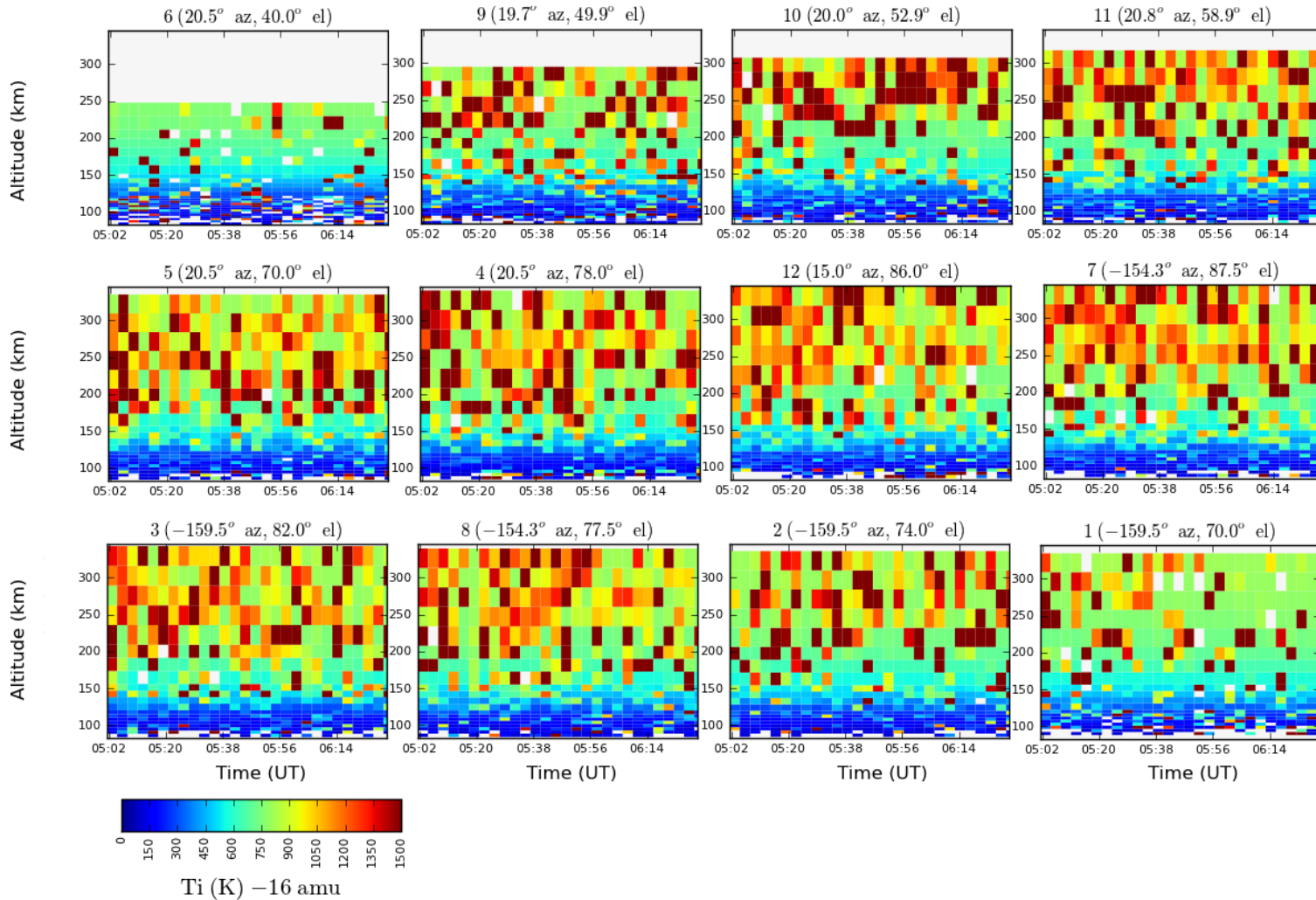
LP 1 min Ne

8-1-2012 5.022 UT - 8-1-2012 6.499 UT



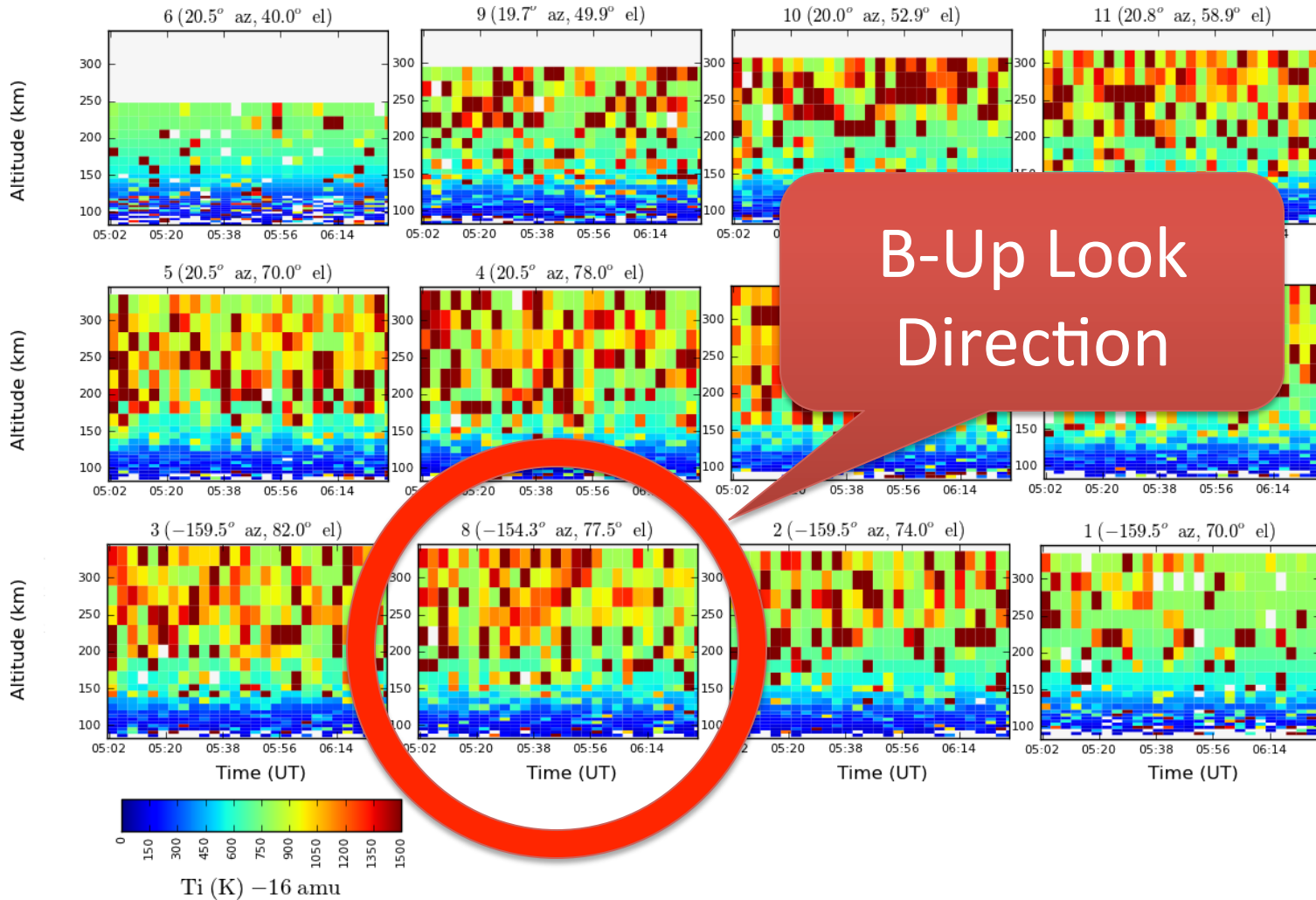
AC 3 min Ti

8-1-2012 5.022 UT - 8-1-2012 6.499 UT

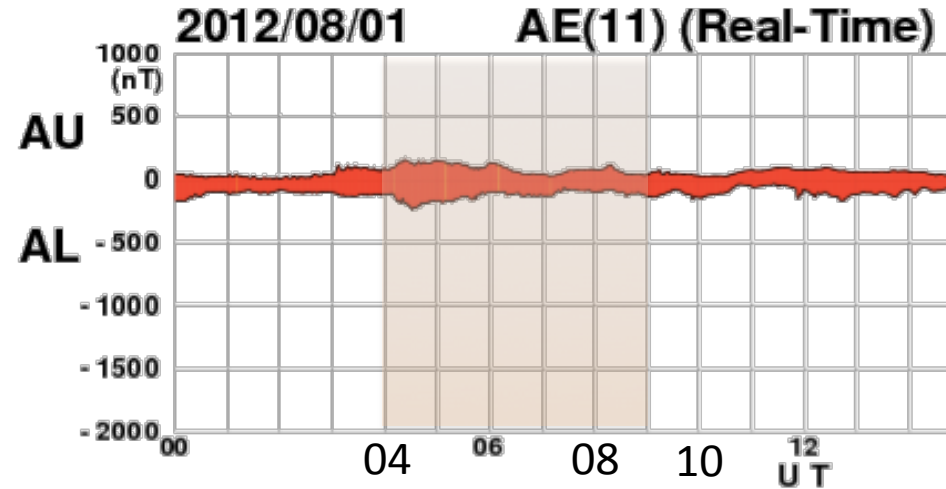
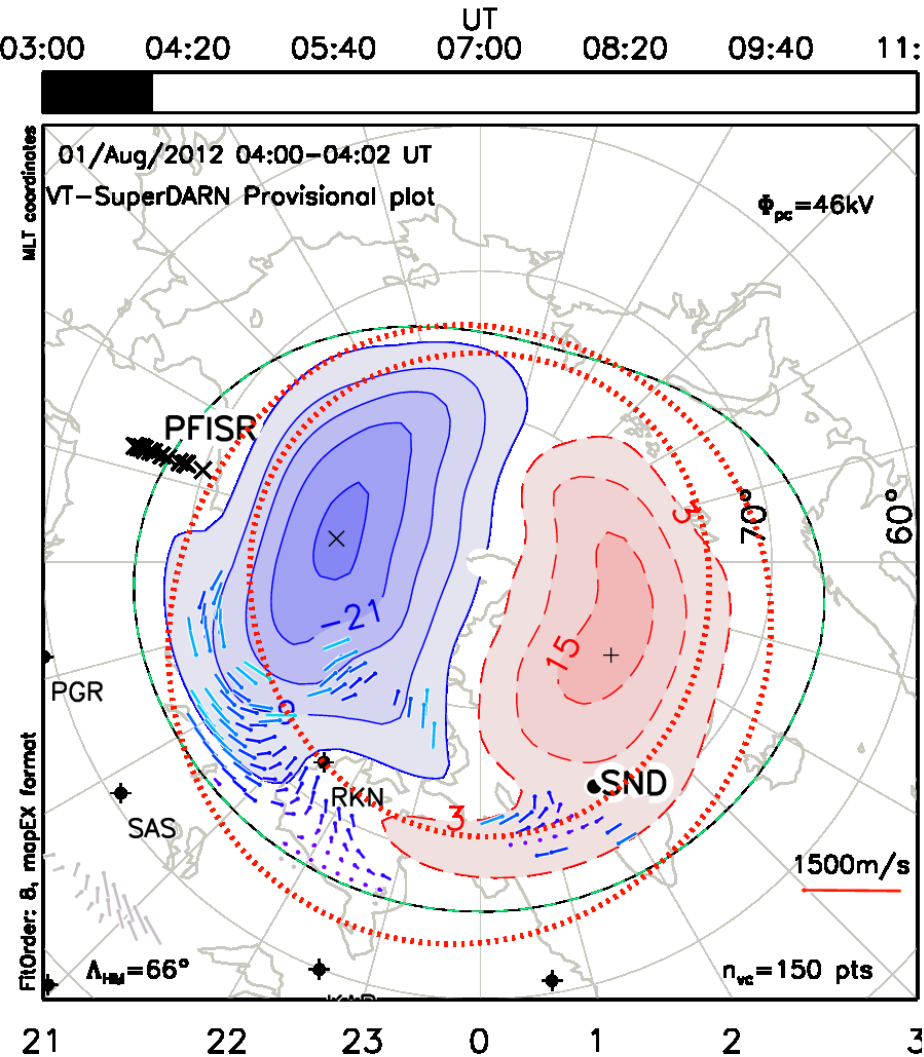


AC 3 min Ti

8-1-2012 5.022 UT - 8-1-2012 6.499 UT



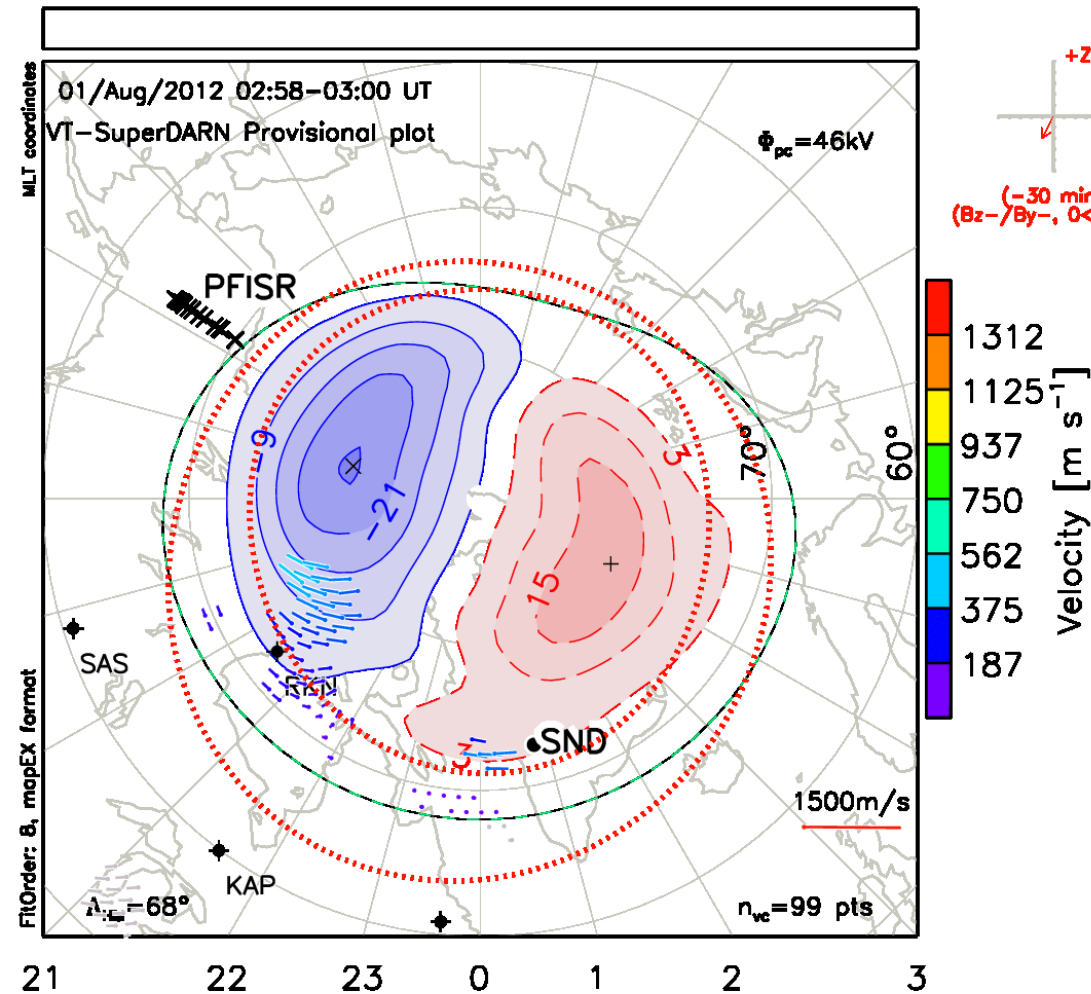
PFISR Altitude: 350–385 km; Felstein Oval for $K_p=1+$



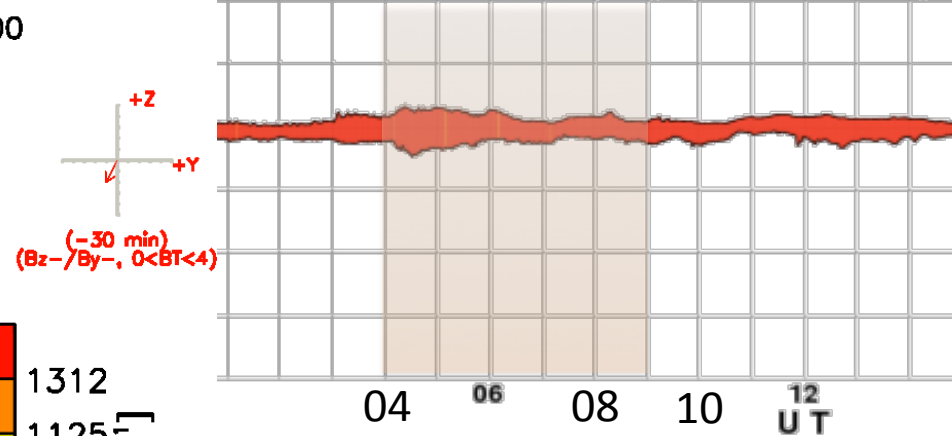
SuperDARN

PFISR Altitude: 350–385 km; Felstein Oval for $K_p=1+$

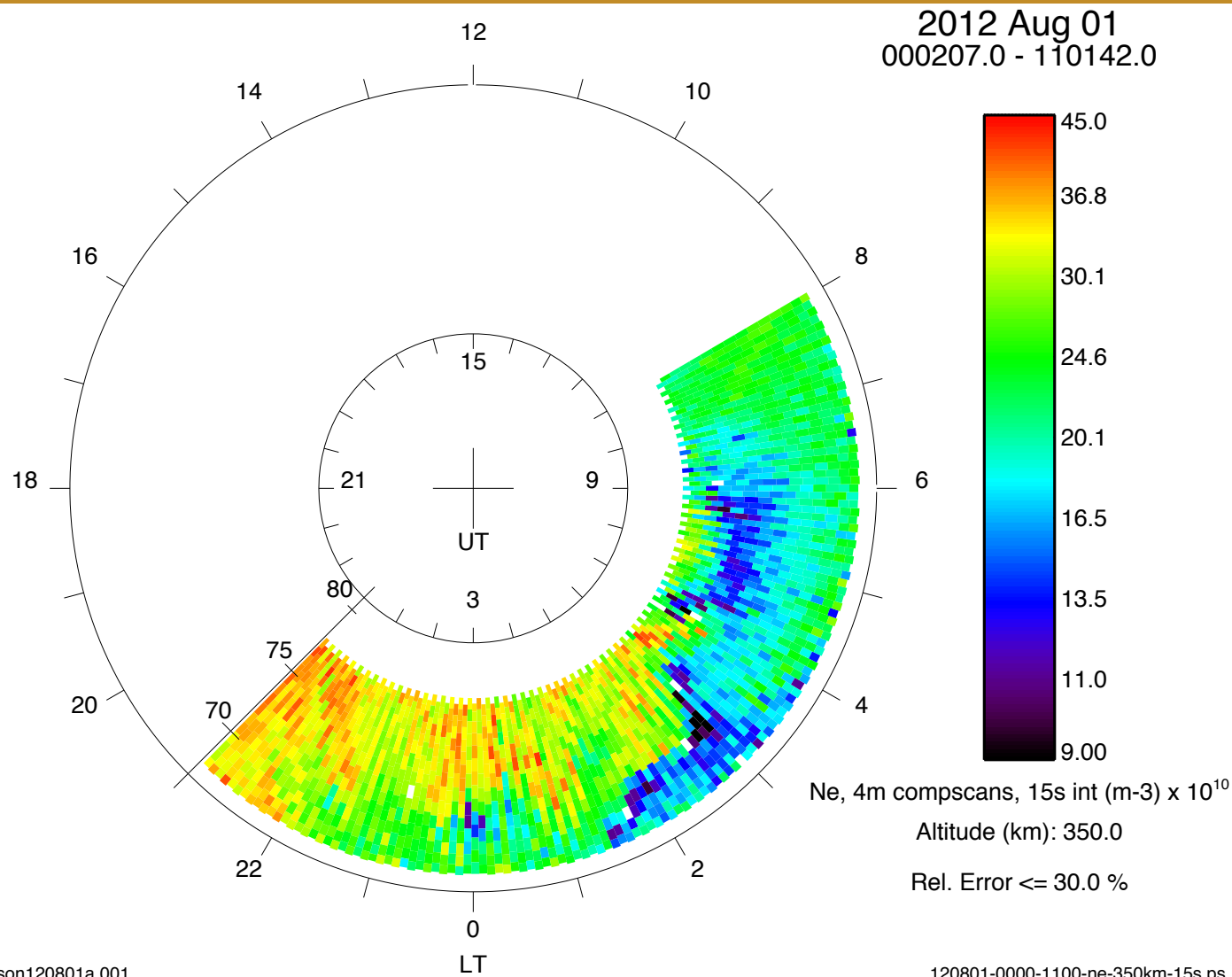
03:00 04:20 05:40 07:00 08:20 09:40 11:00



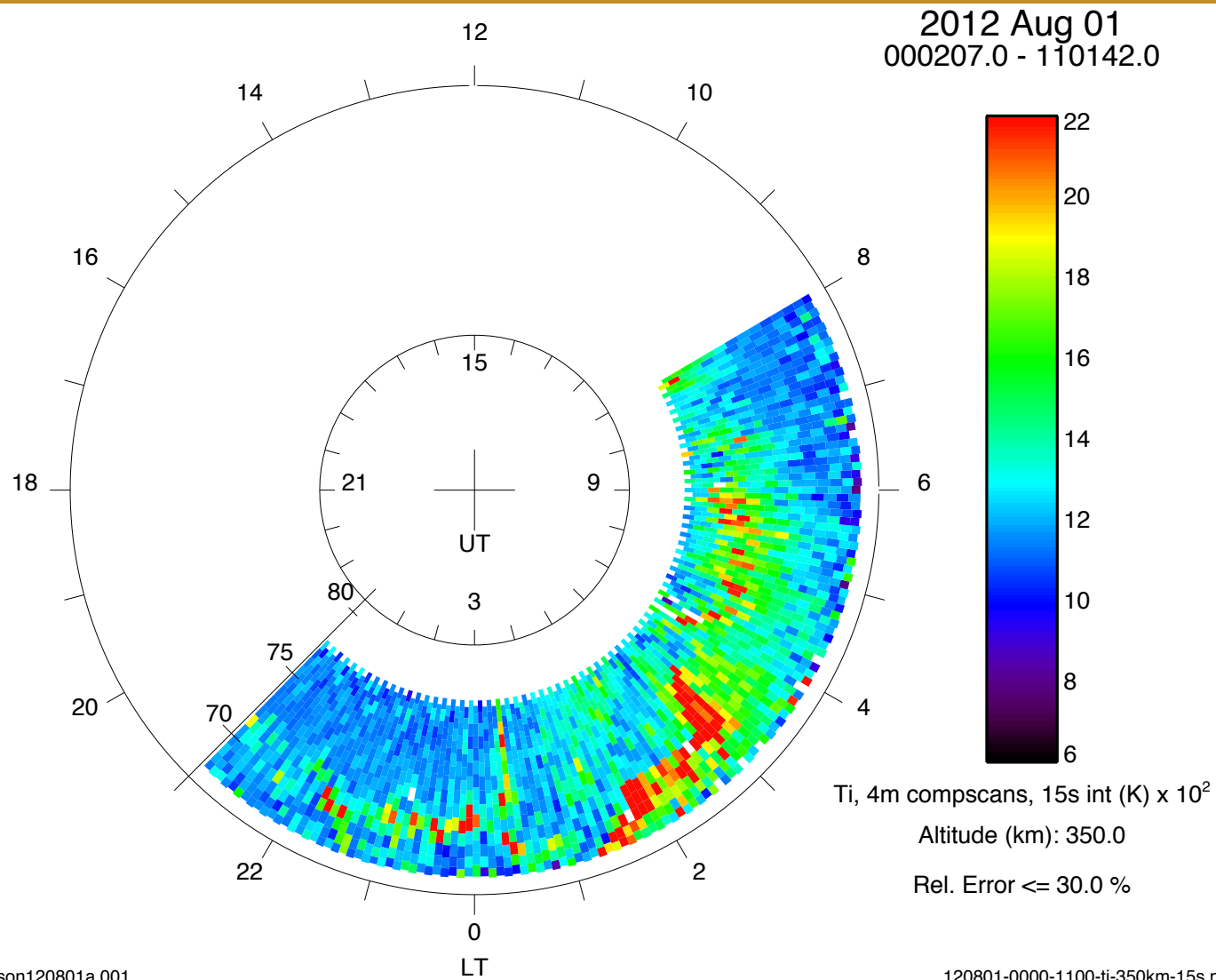
12/08/01 AE(11) (Real-Time)



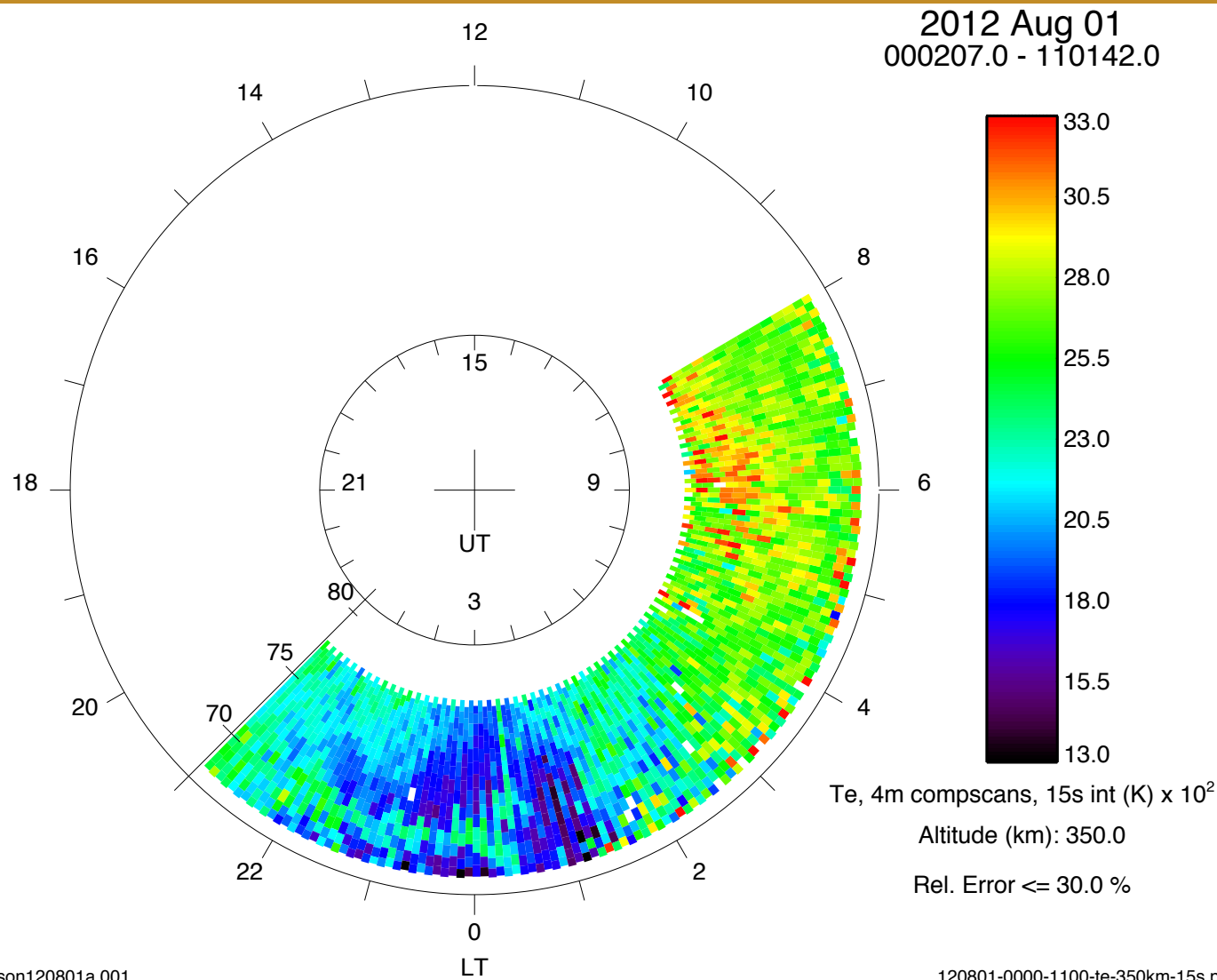
4m Compscans 15 s Ne



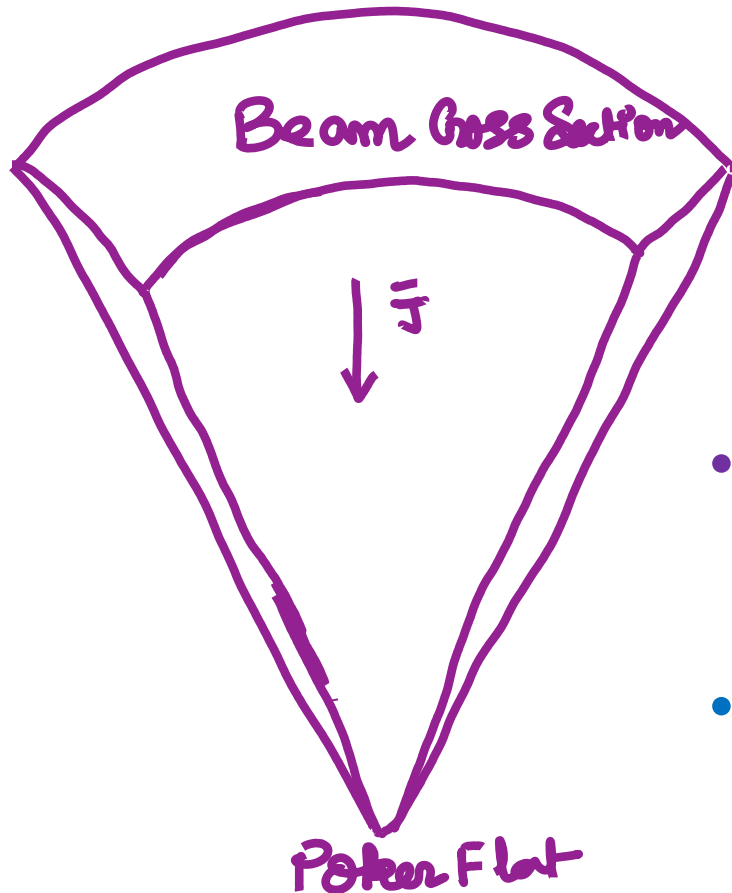
4m Compscans 15 s Ti



4m Compscans 15 s Te



ExB Velocities & Mag. Fields



$$v_e = -\frac{E}{B} v_e / \Omega_e + E \times \Omega_e / B \Omega_e$$

with $V_n = 0, |\Omega_e| > 0$

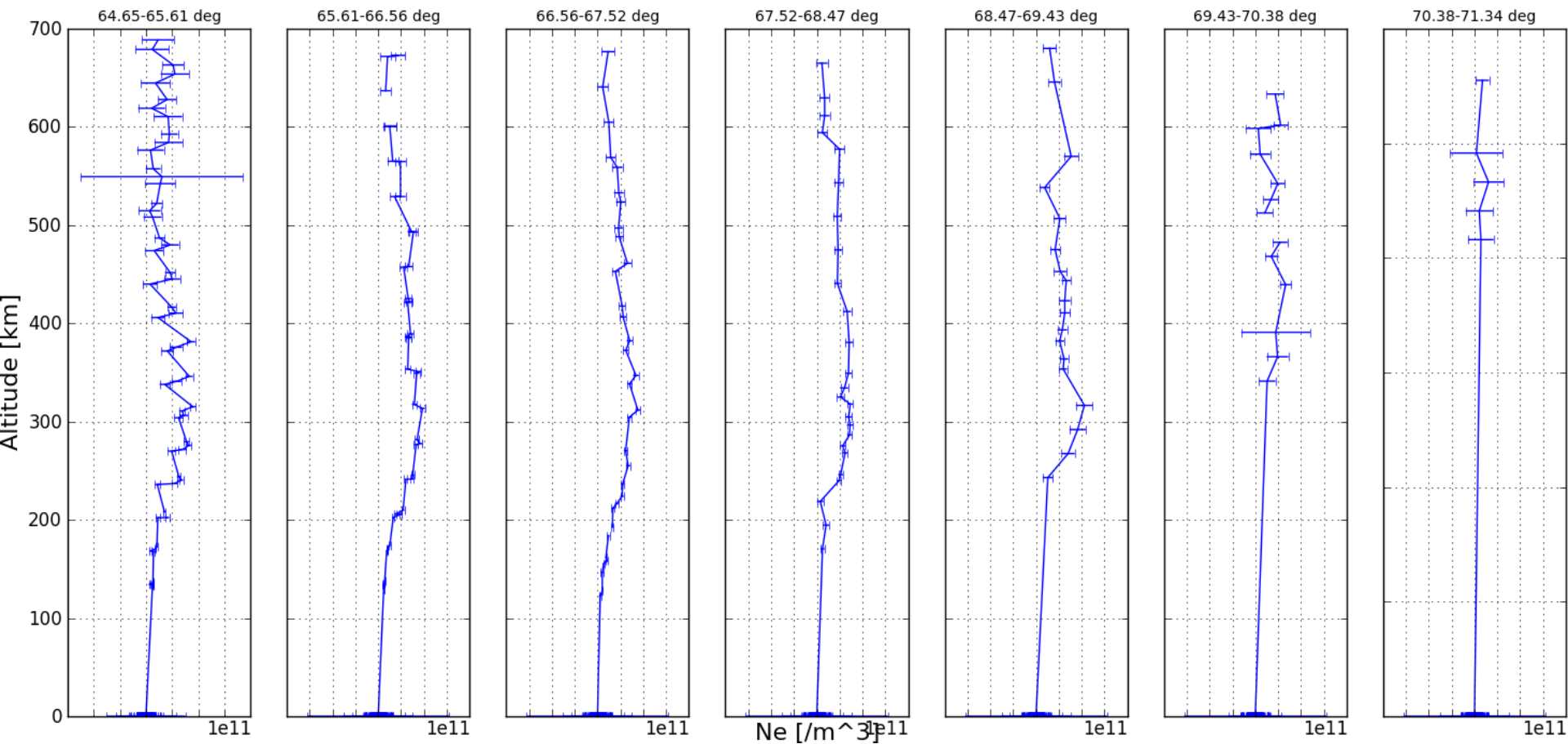
$$v_i = -\frac{E}{B} v_i / \Omega_i + E \times \Omega_i / B \Omega_i$$

with $V_n = 0$

- Rough estimate of velocity from SuperDARN 700m/s
- Magnetic Field Intensity values from various measurements

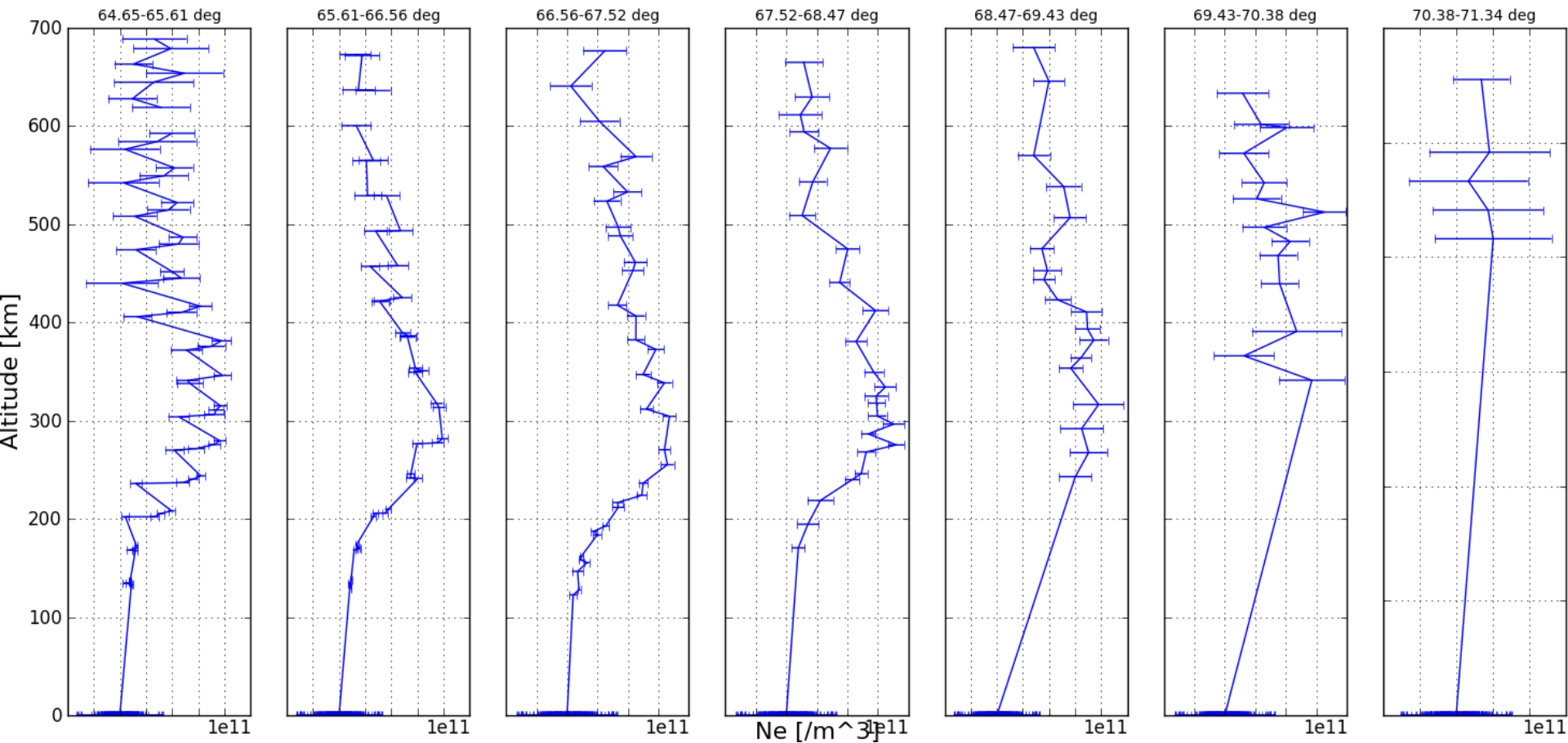
Latitudinal variation of Ne profiles

Plot of Alt. vs Ne, 05:03:11 UT



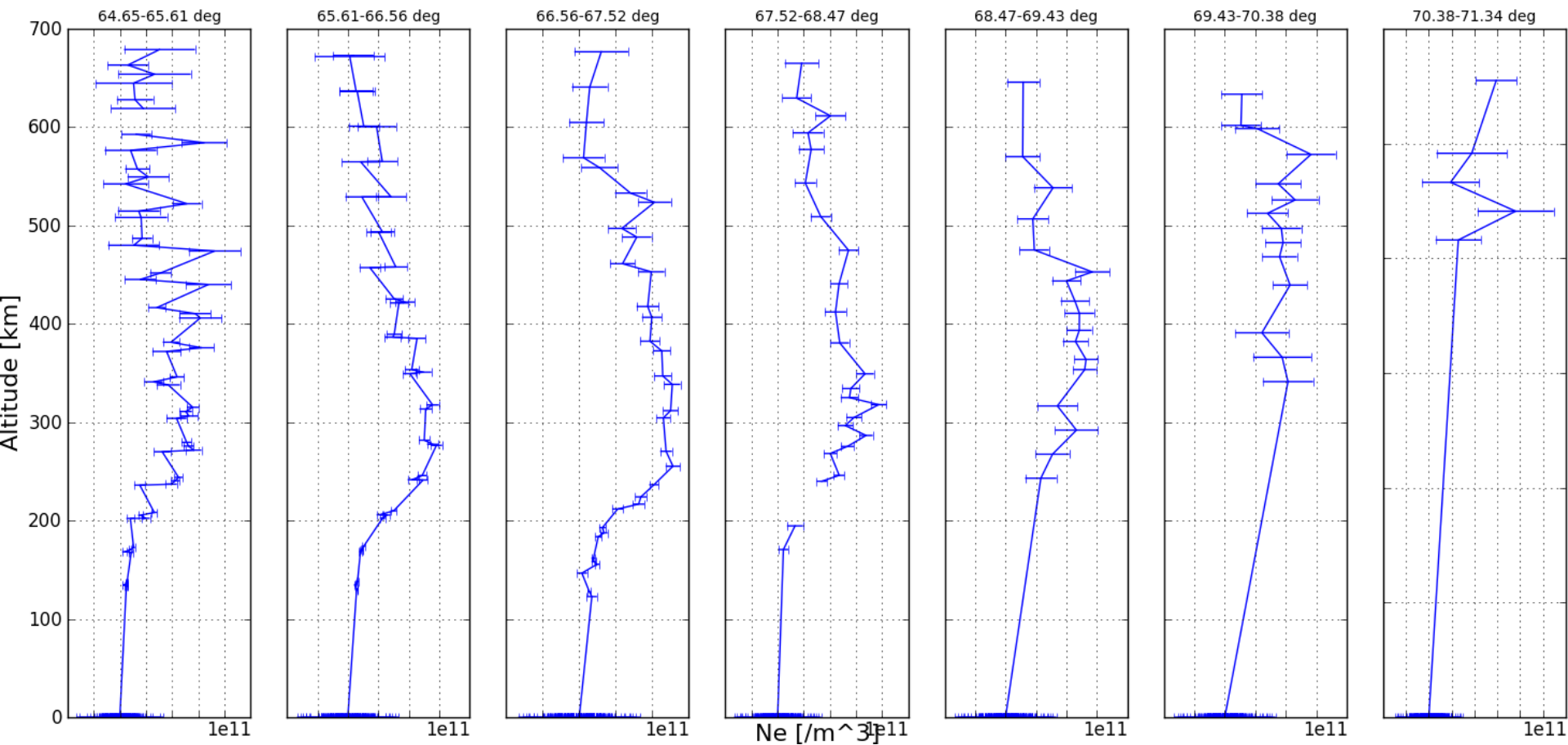
Latitudinal variation of Ne profiles

Plot of Alt. vs Ne, 05:45:37 UT



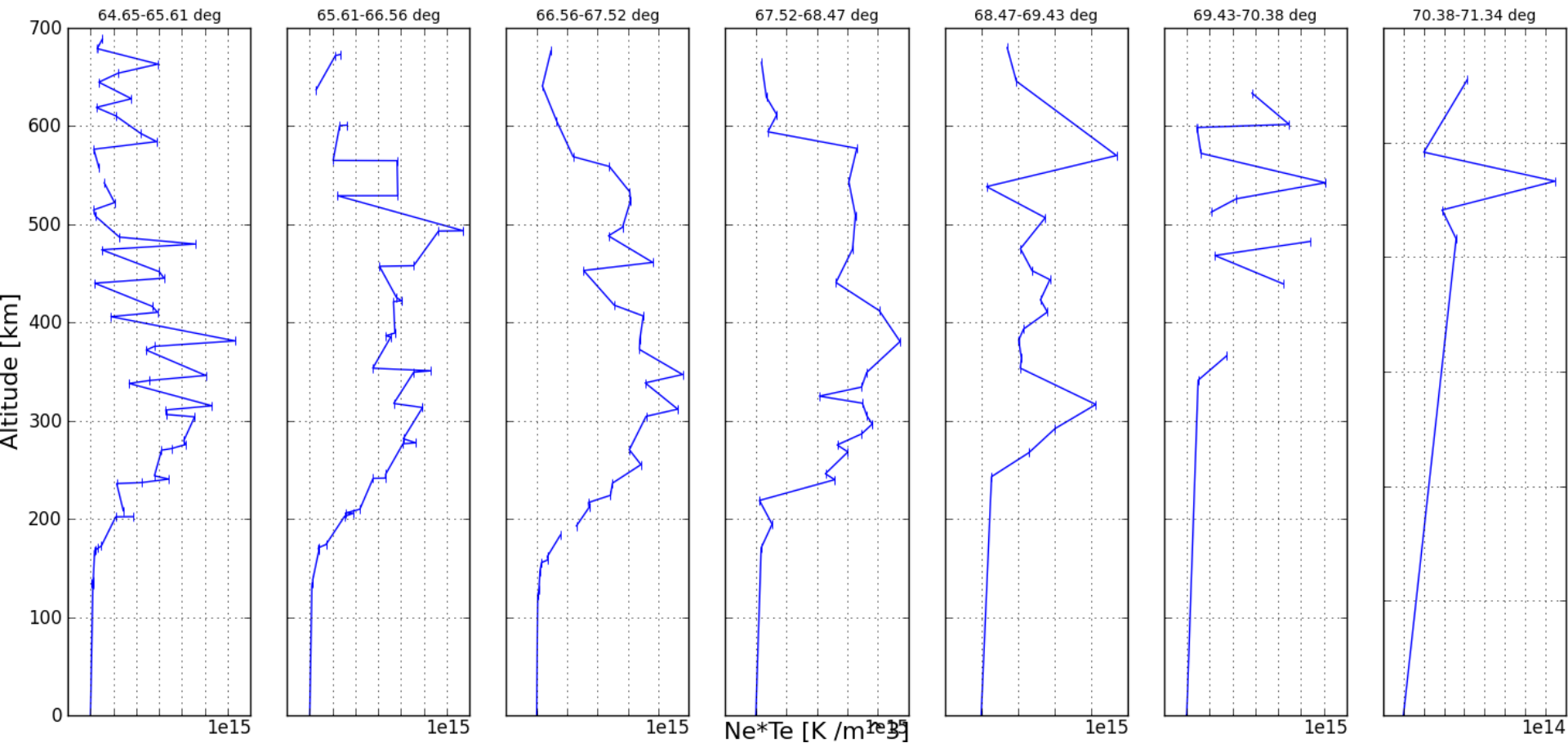
Latitudinal variation of Ne profiles

Plot of Alt. vs Ne, 06:29:18 UT



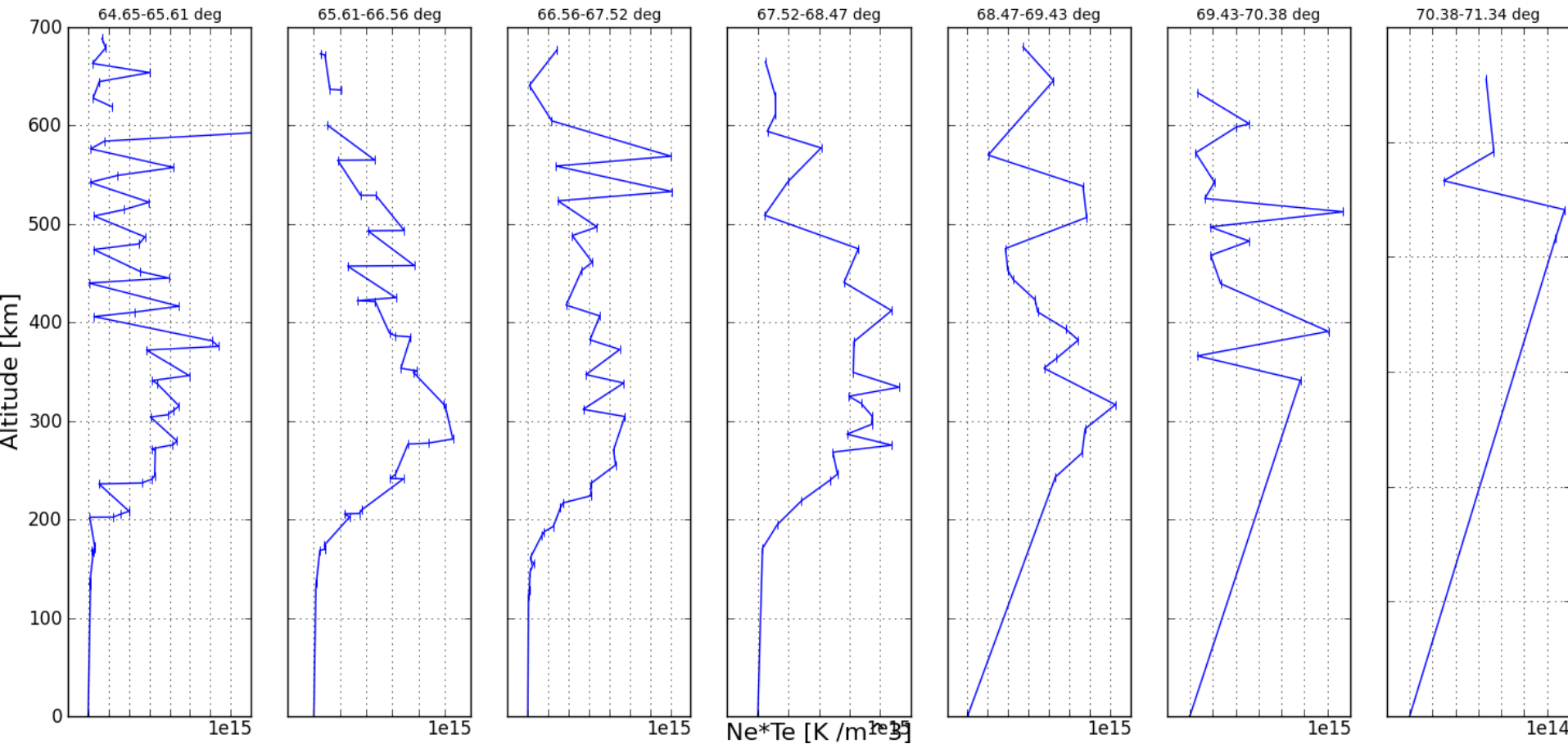
Latitudinal variation of Ne.Te

Plot of Alt. vs Ne*Te, 05:03:11 UT



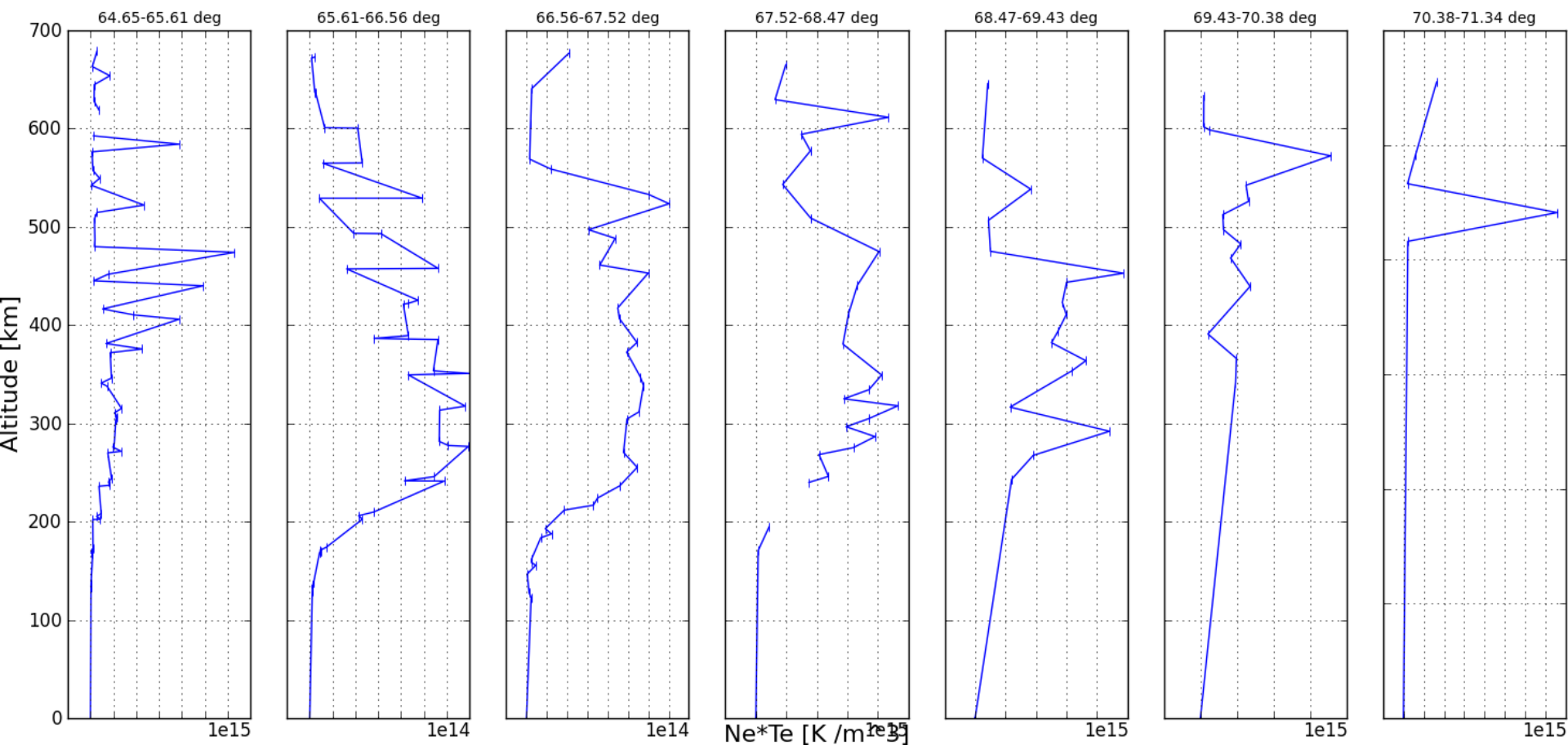
Latitudinal variation of Ne.Te

Plot of Alt. vs Ne*Te, 05:45:37 UT



Latitudinal variation of Ne.Te

Plot of Alt. vs Ne*Te, 06:29:18 UT



Summary and Conclusions

- PFISR campaign conducted on 1AUG2012 0500-0630 UT to study latitudinal structure and dynamics of auroral oval.
- 12 beams located along magnetic meridian used with
 - Long Pulse and Alternating Code modulations
 - 1 and 3 min Integration Times

Summary and Conclusions

- Observations during recovery phase of a small substorm.
 - SW/IMF and geophysical conditions relatively quiet.
 - GMAG shows only significant current system is north of PFISR FOV.
 - PFISR data shows smooth, quiet ionosphere.
- Observations were subauroral, consistent with Feldstein/Holzworth Oval for $K_p = 1+$.
- Slight T_i enhancement observed on **B**-up beam.

Summary and Conclusions

- Observations also made with Sondrestrom ISR.
- At 0400 UT substorm expansion occurs and radar FOV rotates into the auroral zone
- 0400-0900UT
 - Electron depletion
 - Electron and Ion temperature enhancement
- Observations were consistent with Feldstein Oval for $K_p = 1+$.

Summary and Conclusions

- All analysis was done with provisional data.
- Additional error analysis needs to be conducted.
- Sometimes important to look at the same data in different ways.

Acknowledgements

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- Solar Wind and IMF data from ACE is provided by NASA and SWPC.
- AE and Kp index is from the World Data Center in Kyoto.
- Ground magnetometer data from University of Alaska Fairbanks.
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- SuperDARN data funded by NSF, CSA, and NSERC provided J.M. Ruohoniemi, K. McWilliams, and J.P. St. Maurice.

References

Akasofu, S. I. The development of the auroral substorm. *Planetary and Space Science*, 12(4):273–282, 1964.

Holzworth, R. H. and C.-I. Meng (1975), Mathematical representation of the auroral oval, *Geophys. Res. Lett.*, 2 (9), 377–380, doi:10.1029/GL002i009p00377.

Thank you!



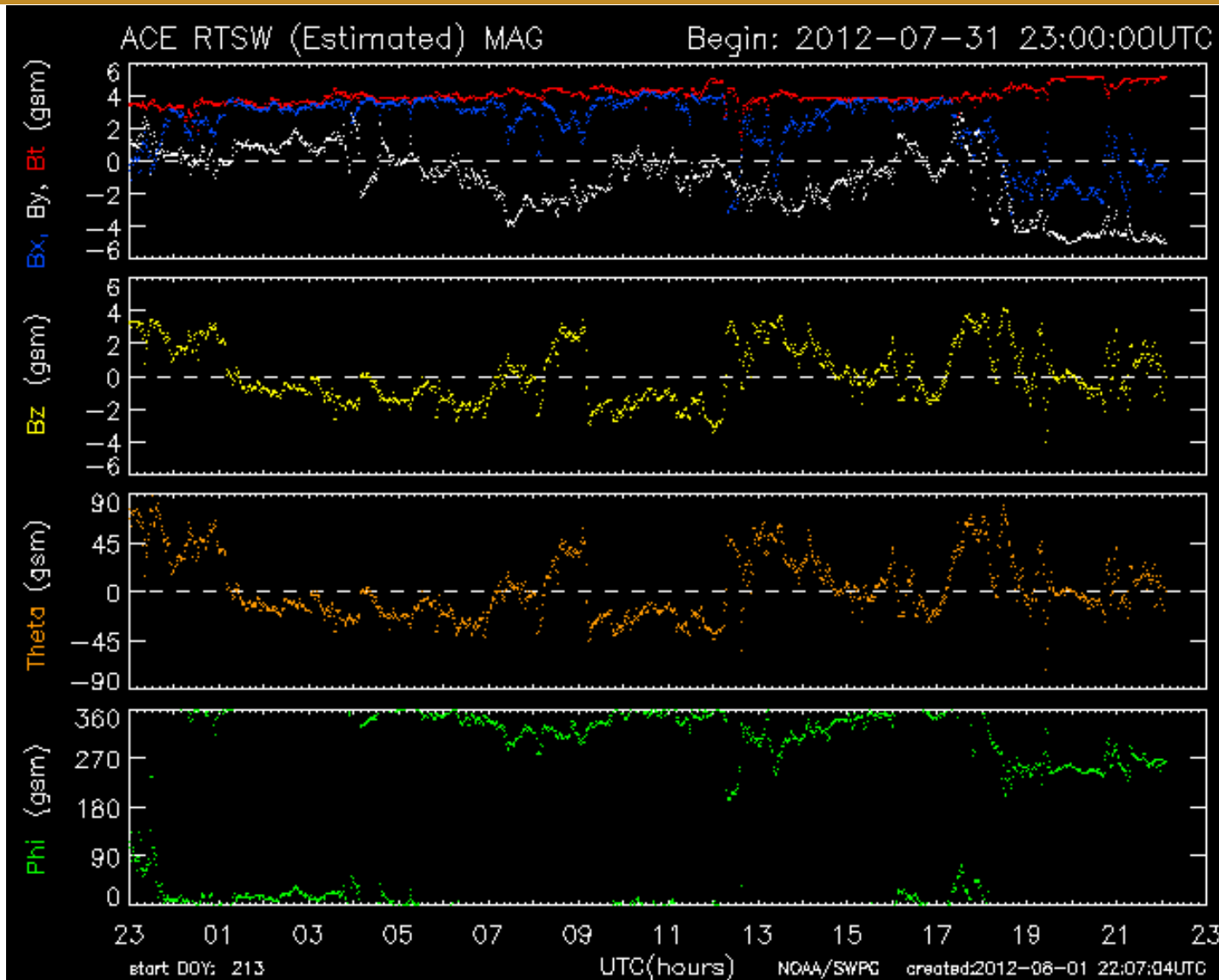
Student exercise/competition



Move together, people! Show a little collective behavior!

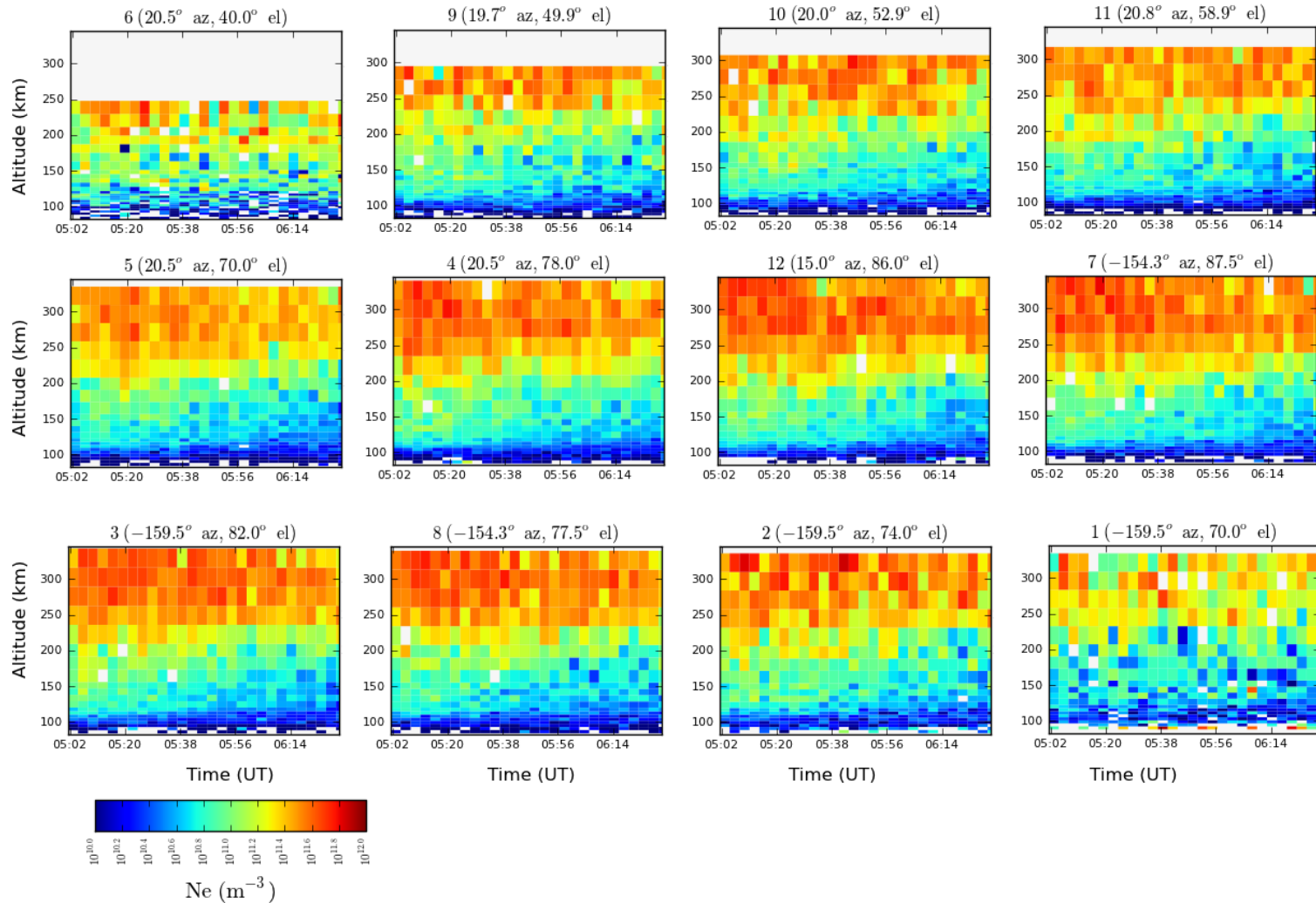
Backup Slides

ACE SW



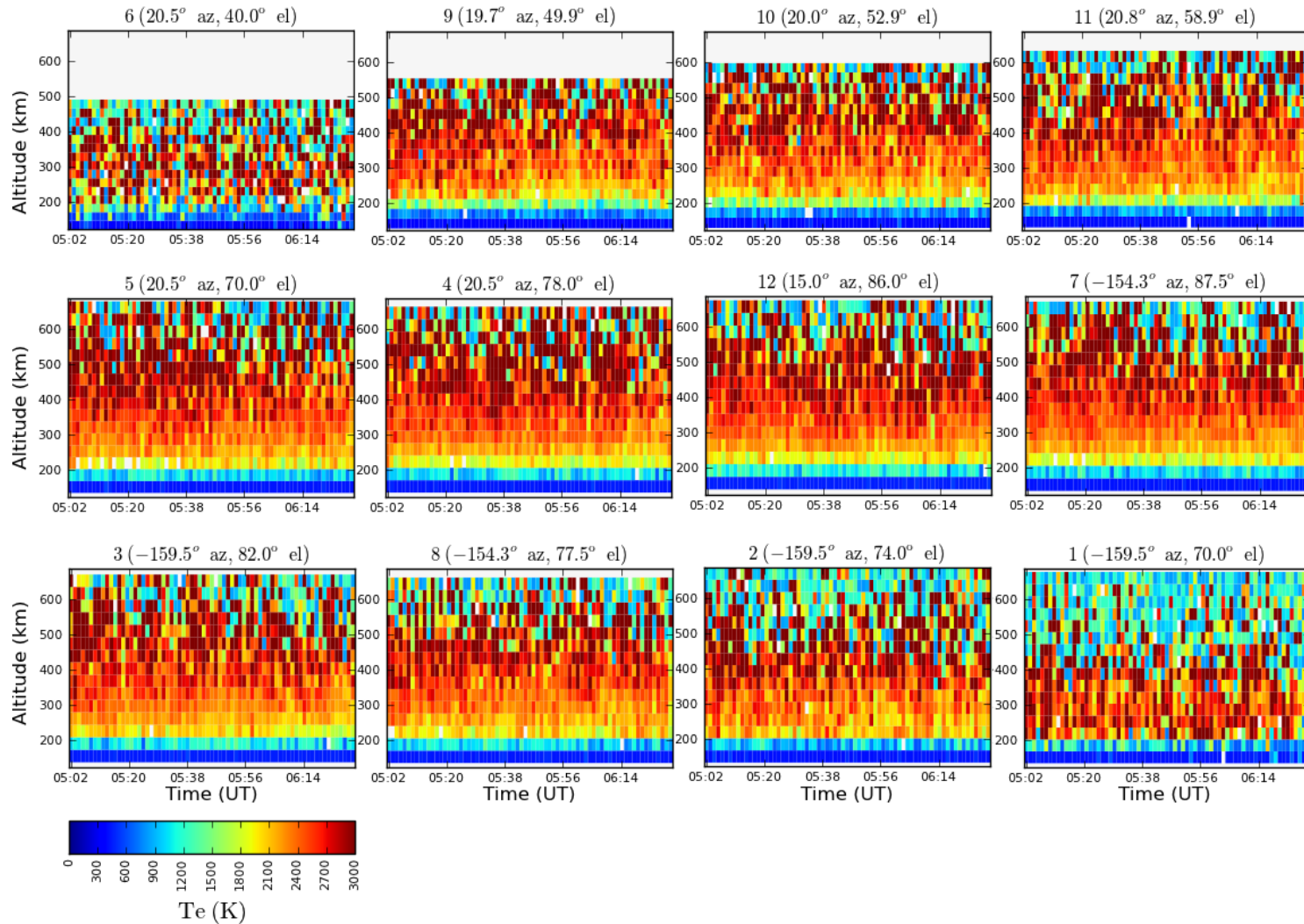
AC 3 min Ne

8-1-2012 5.022 UT - 8-1-2012 6.499 UT

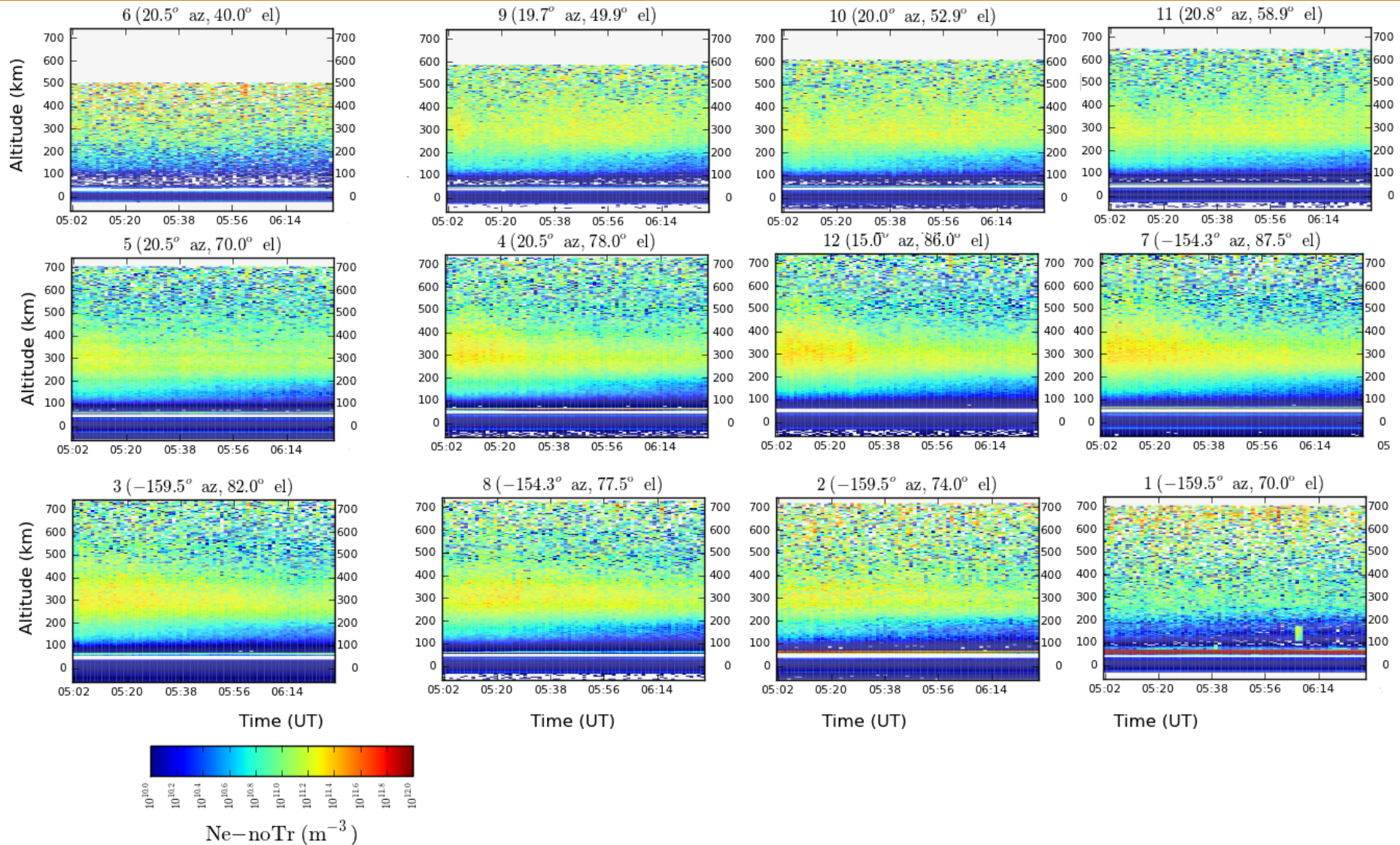


LP 1 min Te

8-1-2012 5.022 UT - 8-1-2012 6.499 UT

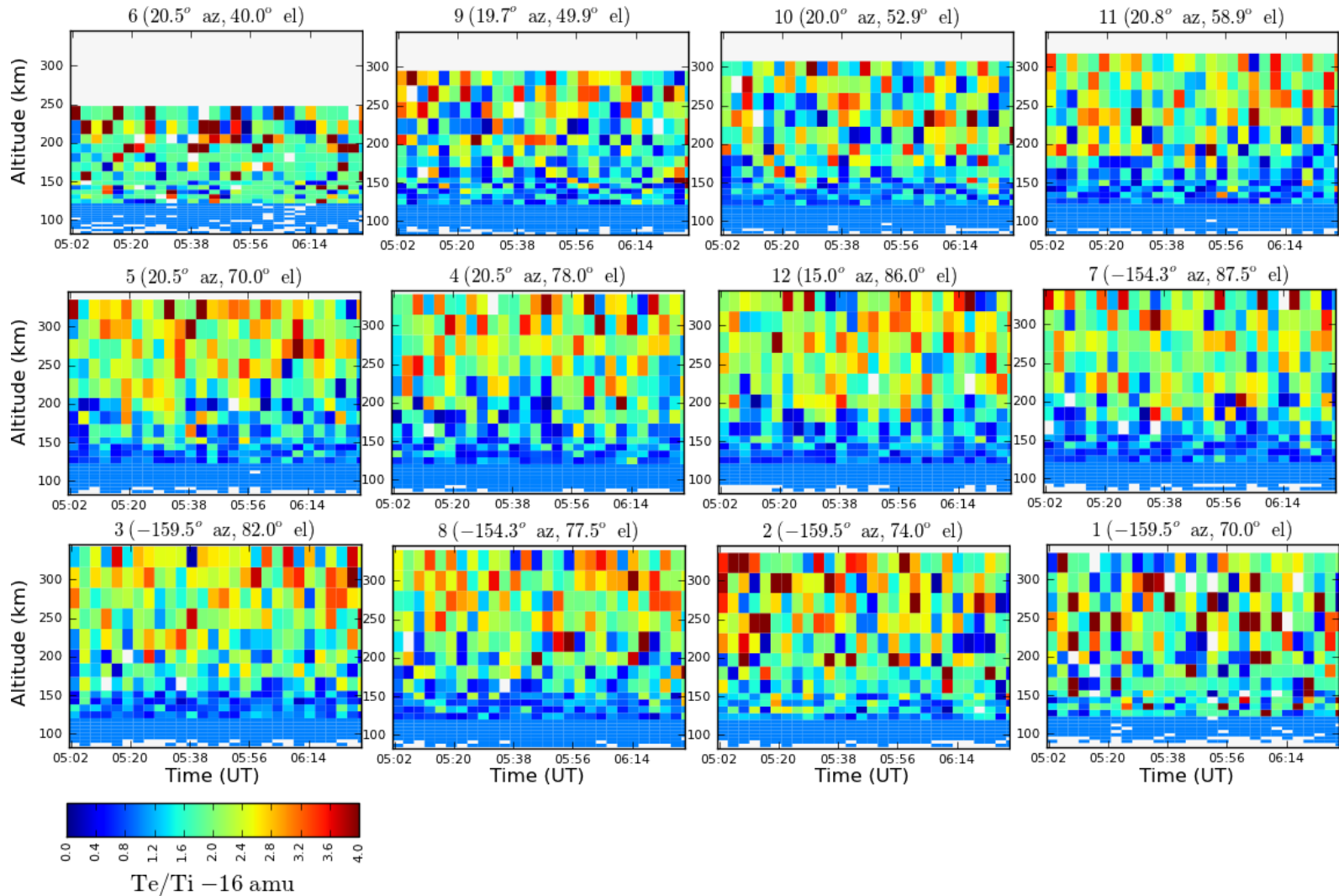


LP 1 min Ne (Not Corrected)



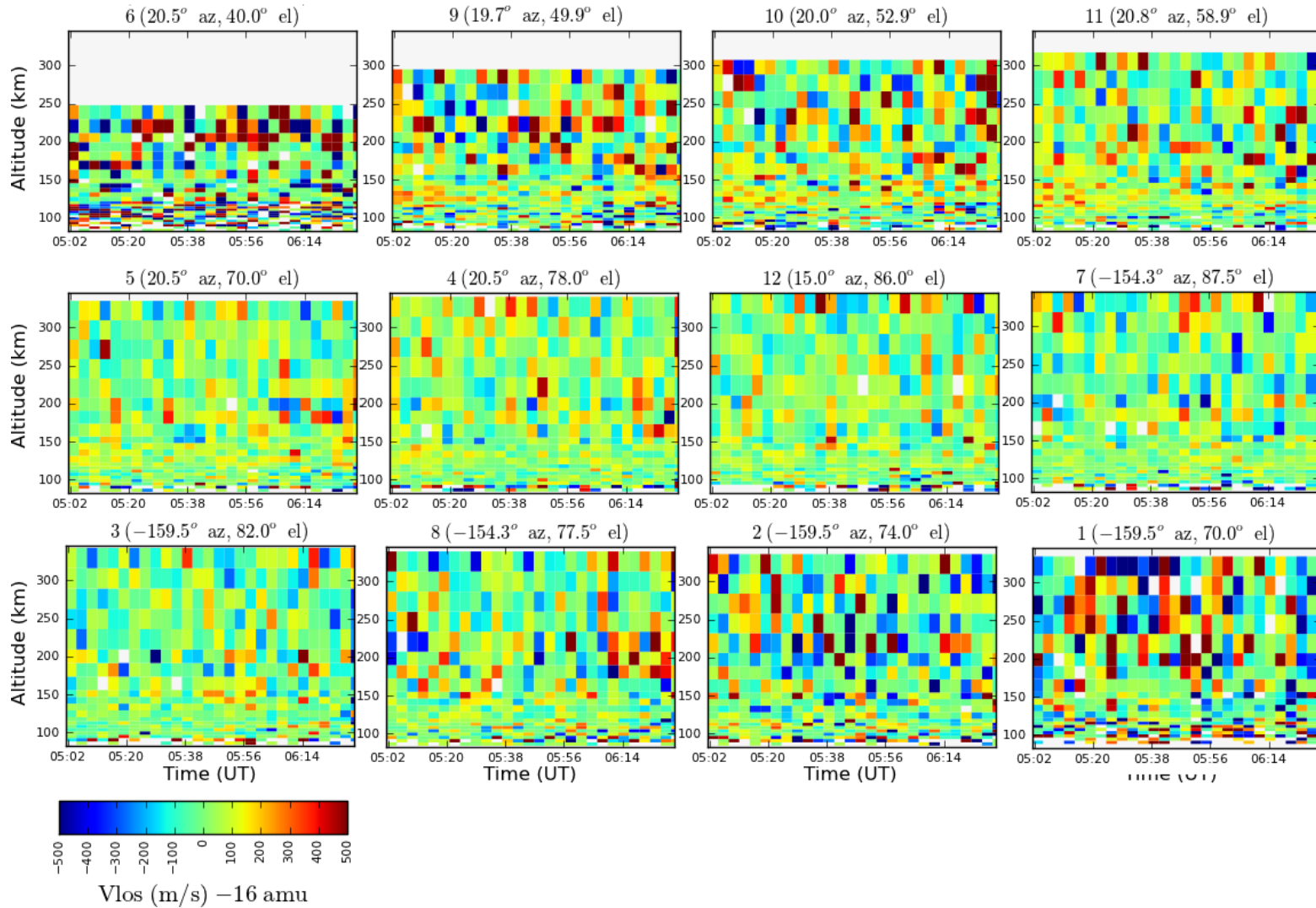
AC 3 min Te/Ti

8-1-2012 5.022 UT - 8-1-2012 6.499 UT

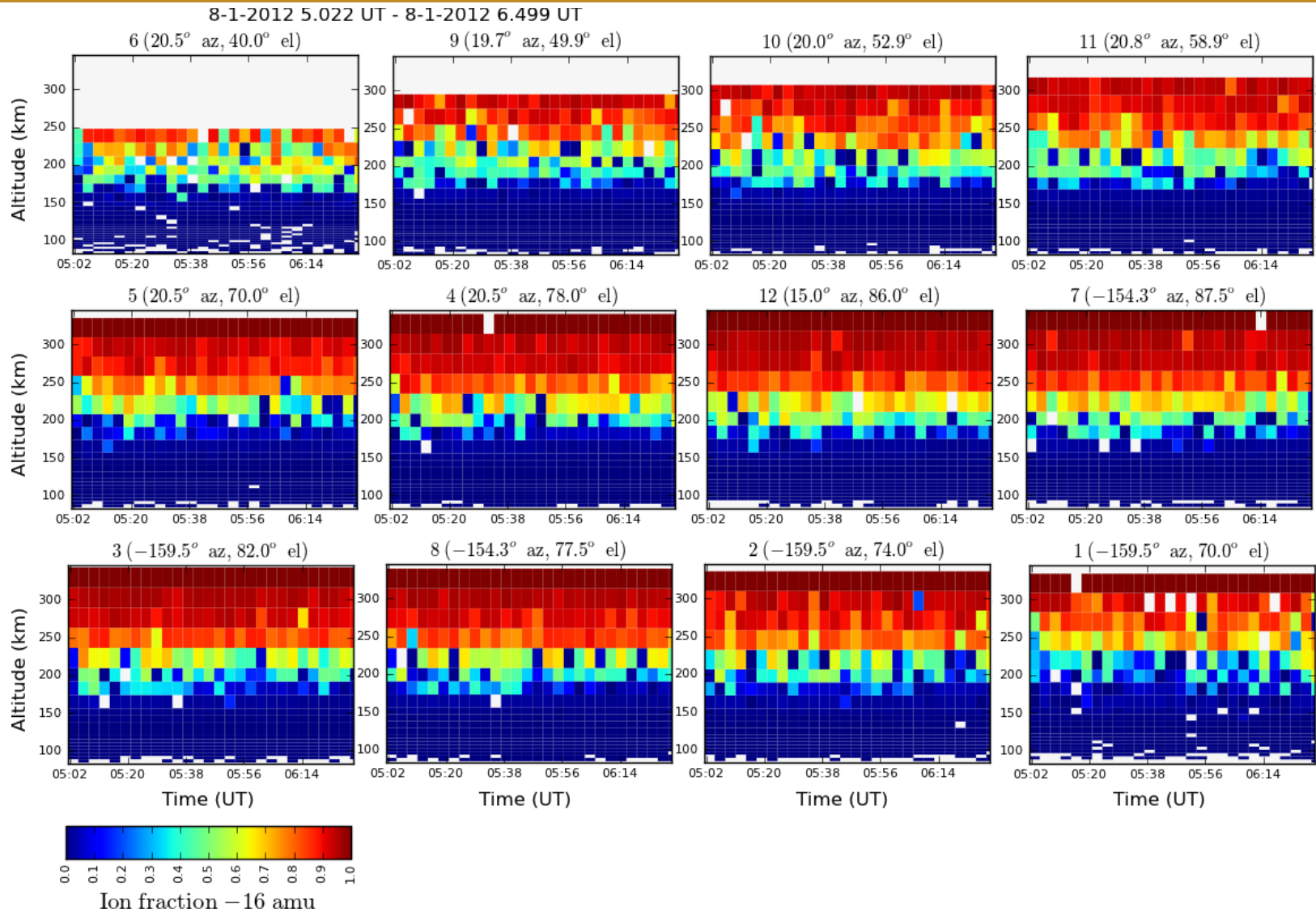


AC 3 min LOS Velocity

8-1-2012 5.022 UT - 8-1-2012 6.499 UT



AC 3 min O+ Fraction



AC 3 min ν_{in}

8-1-2012 5.022 UT - 8-1-2012 6.499 UT

